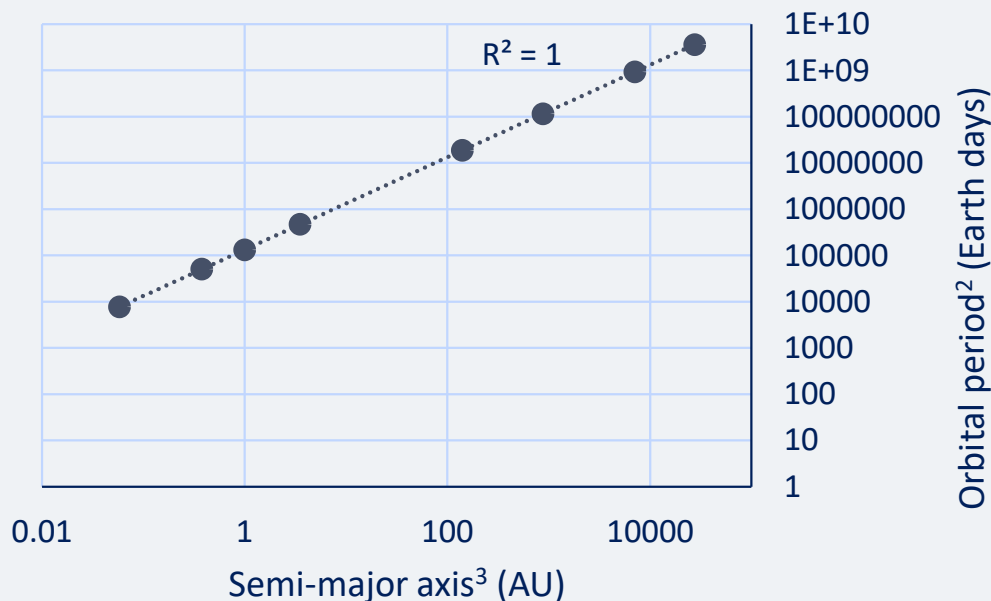




TASK 1: INVESTIGATE KEPLER'S 3RD LAW



- The data points plot in a straight line.
- The line of best fit has an R^2 value of 1. This tells us there is a strong relationship between the 2 variables.
- The squares of the orbital periods are directly proportional to the cubes of the semi-major axis.

TASK 2: APPLY KEPLER'S 3RD LAW

$$T^2 = \frac{4\pi^2}{GM} a^3$$

T = orbital period

a = semi-major axis

G = gravitational constant ($G = 6.67 \times 10^{-11} \text{ m}^3 \text{ kg}^{-1} \text{ s}^{-2}$)

M = mass of the Sun

Rearrange the equation to make M the subject:

Step 1

Divide both sides by a^3

$$\frac{T^2}{a^3} = \frac{4\pi^2}{GM}$$

Step 2

Multiply both sides by M

$$M \left(\frac{T^2}{a^3} \right) = \frac{4\pi^2}{G}$$

Step 3

Divide both sides by $\left(\frac{T^2}{a^3} \right)$

$$M = \frac{4\pi^2 a^3}{G T^2}$$

Work out the value of M . Express your value to 3 significant figures, using scientific notation:

	Mercury	Venus	Earth	Mars	Jupiter	Saturn	Uranus	Neptune
Calculated value of M (kg)	1.99×10^{30}	1.98×10^{30}	1.99×10^{30}	1.98×10^{30}	1.98×10^{30}	2.01×10^{30}	1.99×10^{30}	2.01×10^{30}

The true mass of the Sun is $1.989 \times 10^{30} \text{ kg}$.

All your values for M should be relatively similar and close to this value